

Most of the energy on Earth comes to us from the Sun.

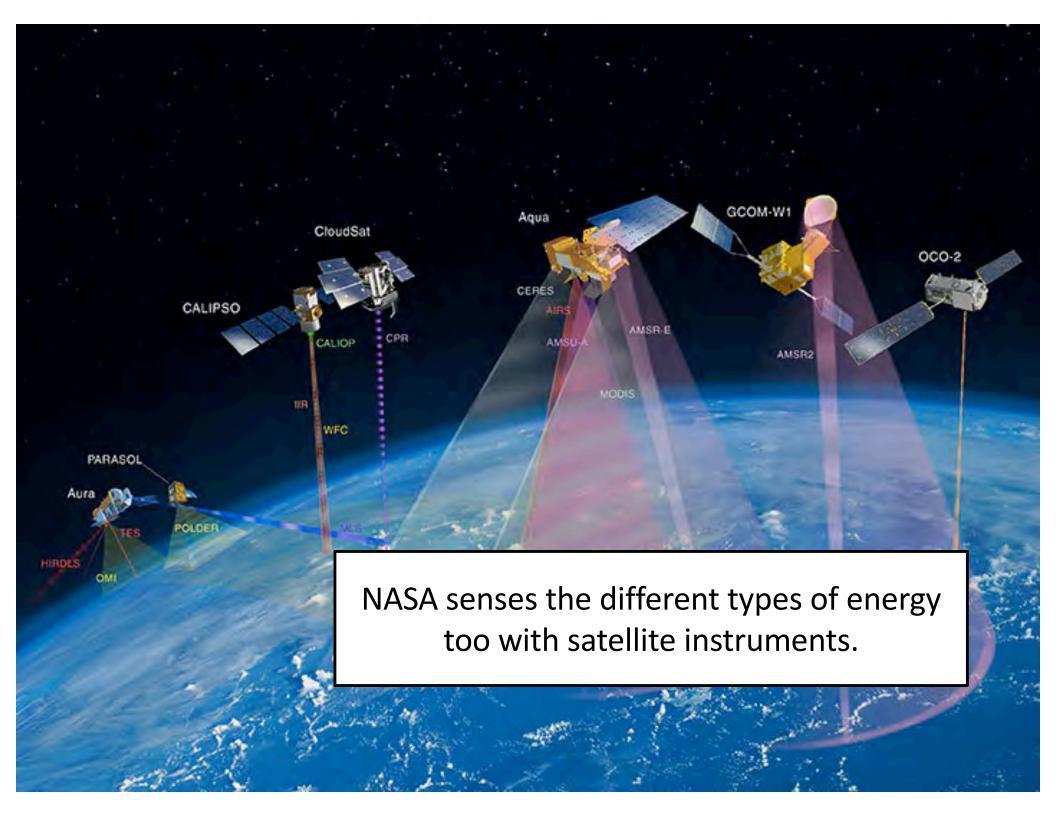
Did you know?: The amount of sunlight that reaches the Earth is equal to approximately 6 60w light bulbs for every square meter of the surface.

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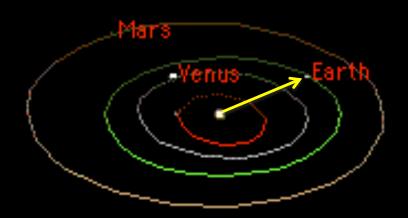


We can sense that energy in different ways. We see the things around us because of visible light...





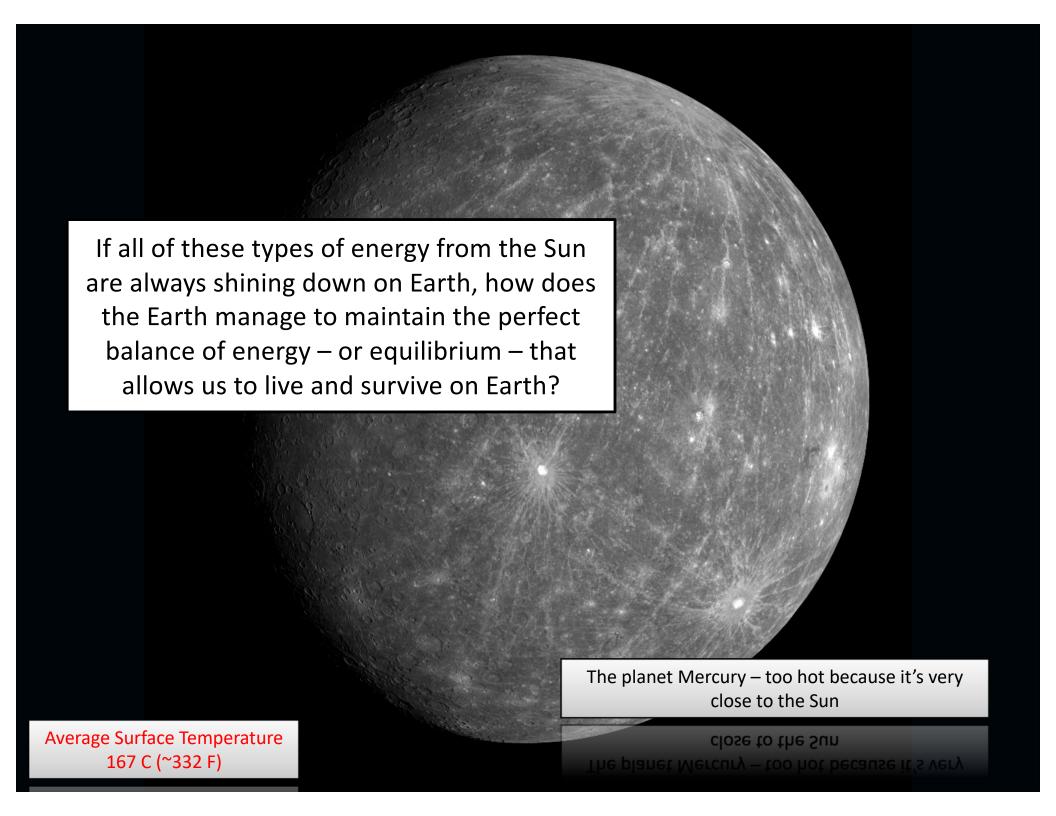
If all of these types of energy from the Sun are always shining down on Earth, how does the Earth manage to maintain the perfect balance of energy – or equilibrium – that allows us to live and survive on Earth?



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The Sun – hot though it is - is a tiny part of Earth's environment. The rest is cold, dark space.

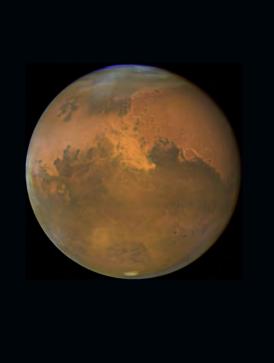
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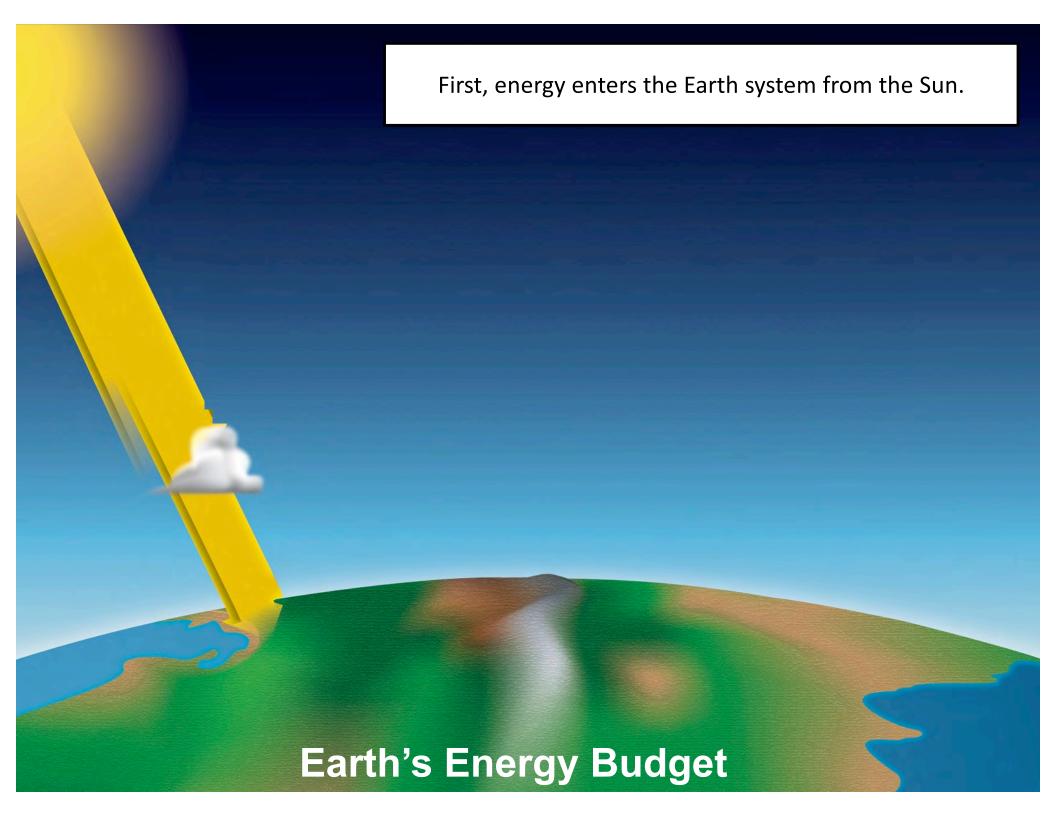


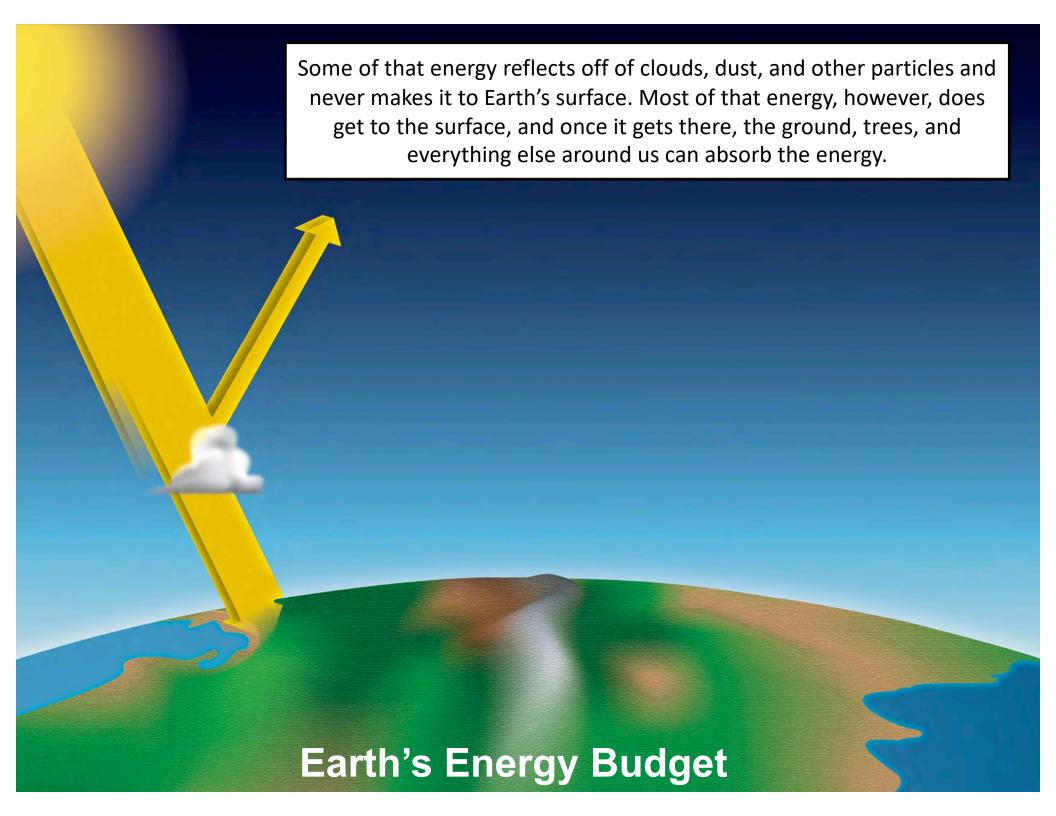


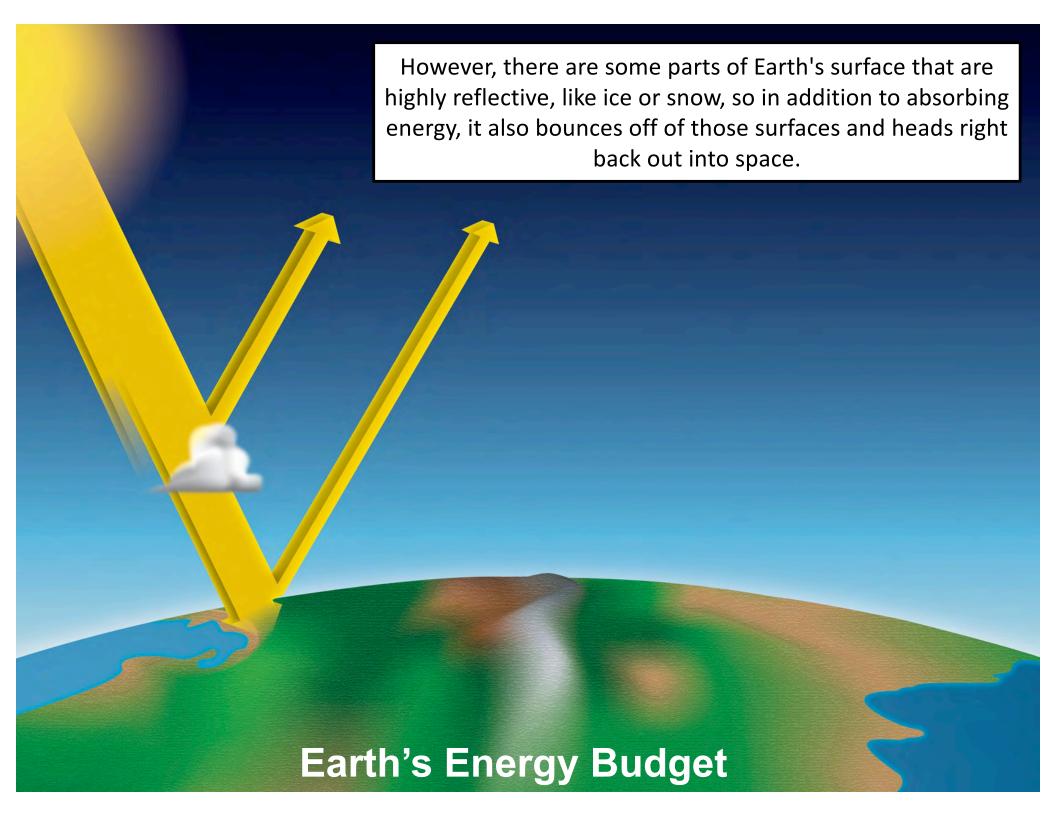




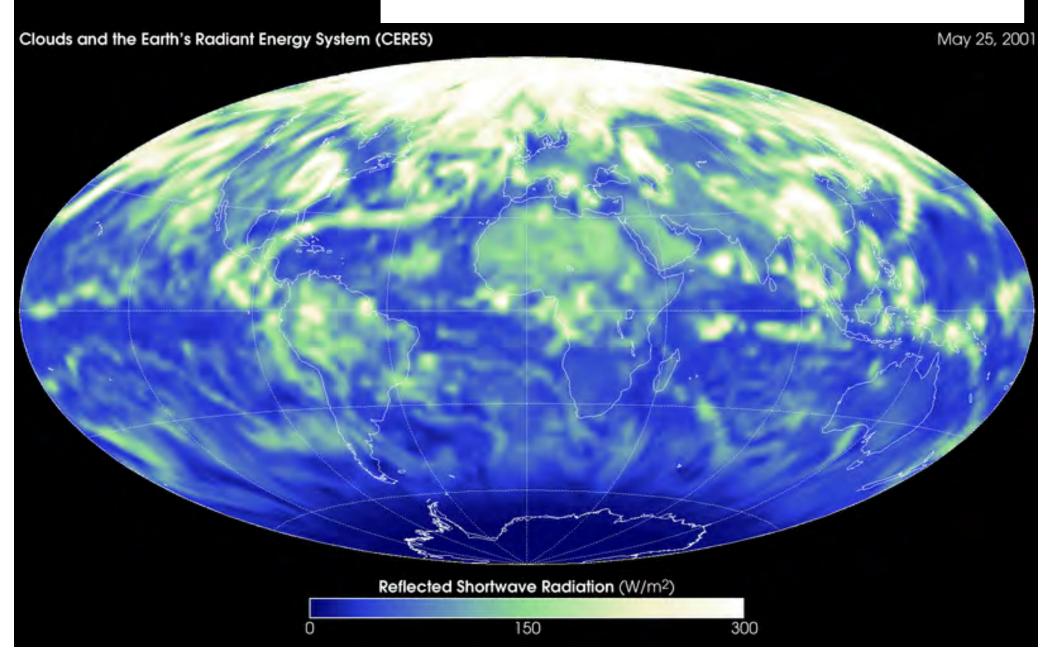
The temperatures of Earth and all the planets are determined by their "Energy Budget."

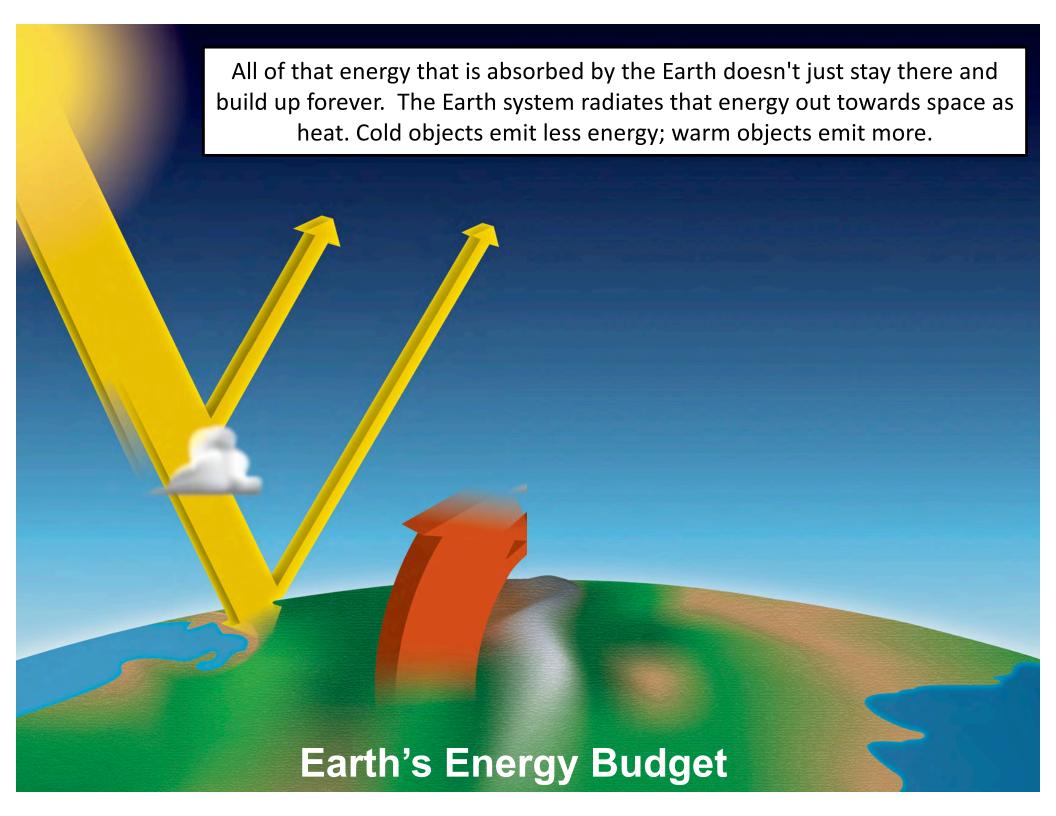


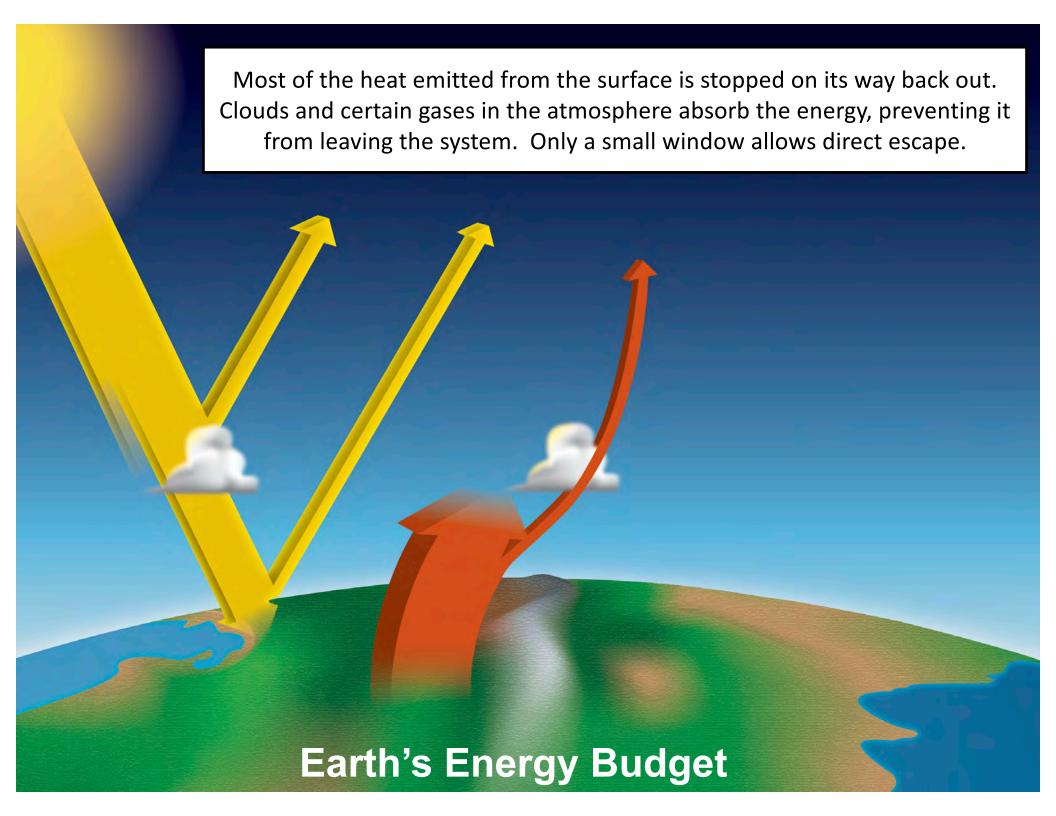


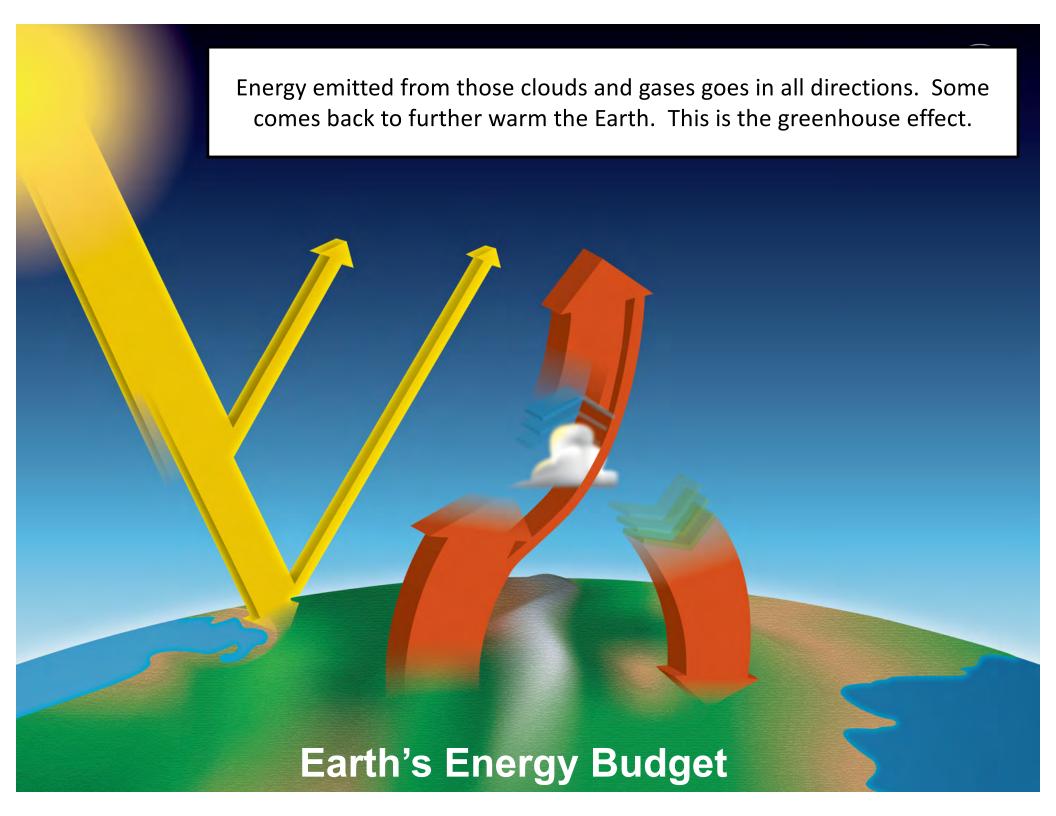


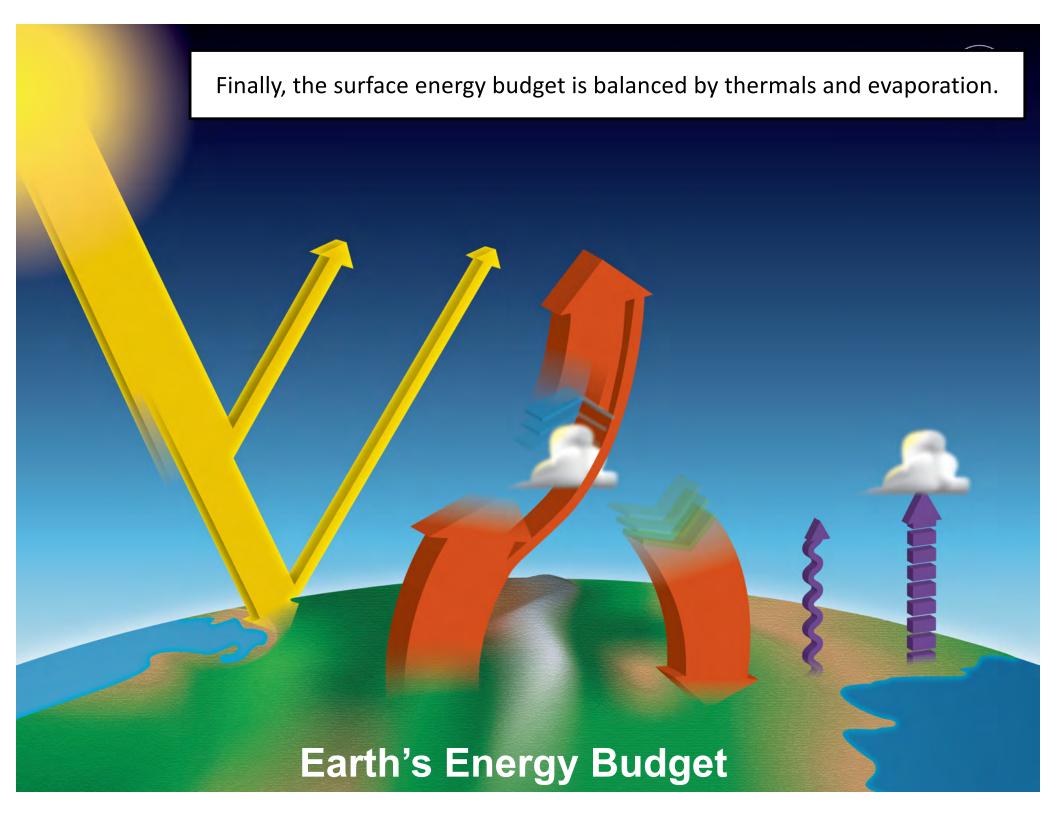
The world in reflected sunlight, May 25, 2001. Clouds, deserts and Arctic ice are bright. The south pole is in winter darkness with no sunlight to reflect.

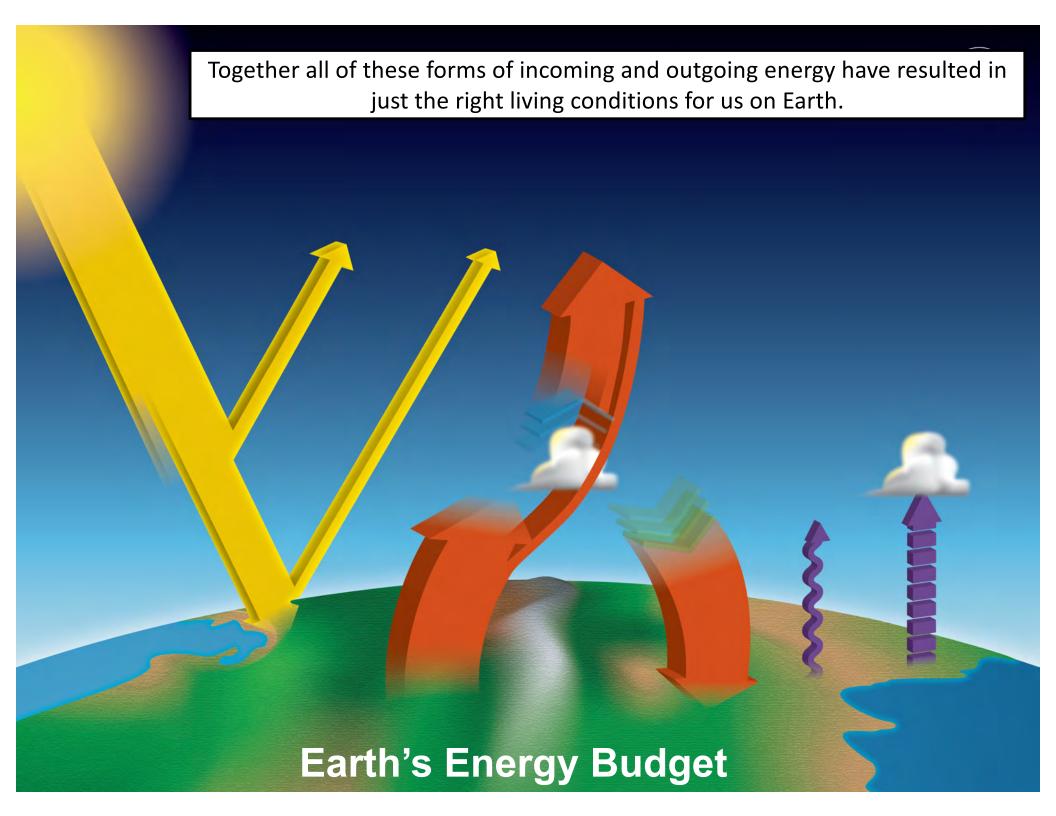




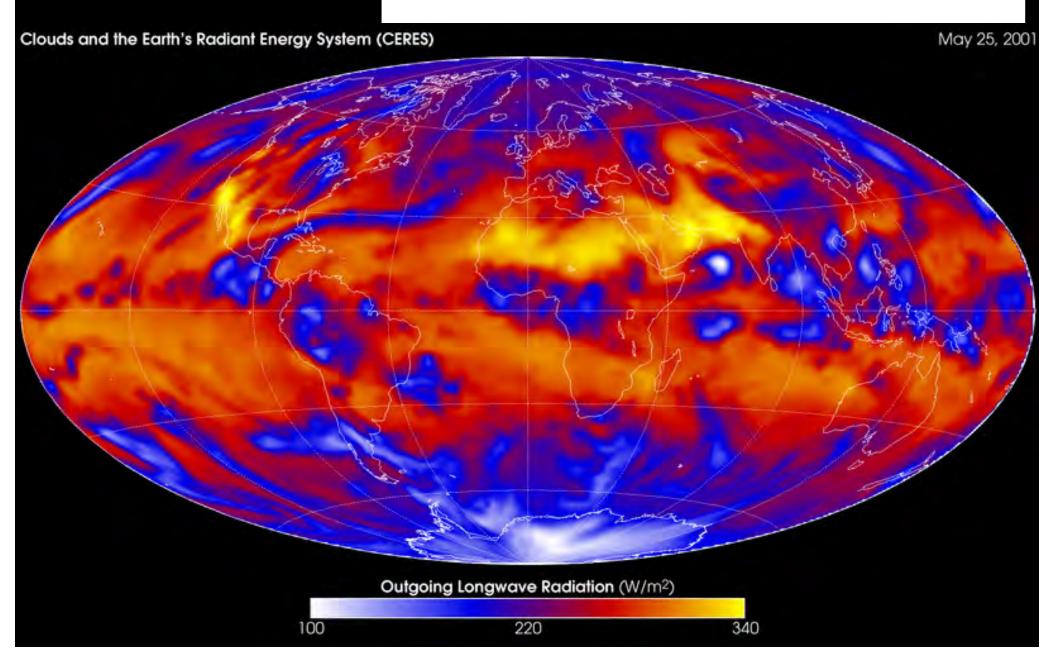


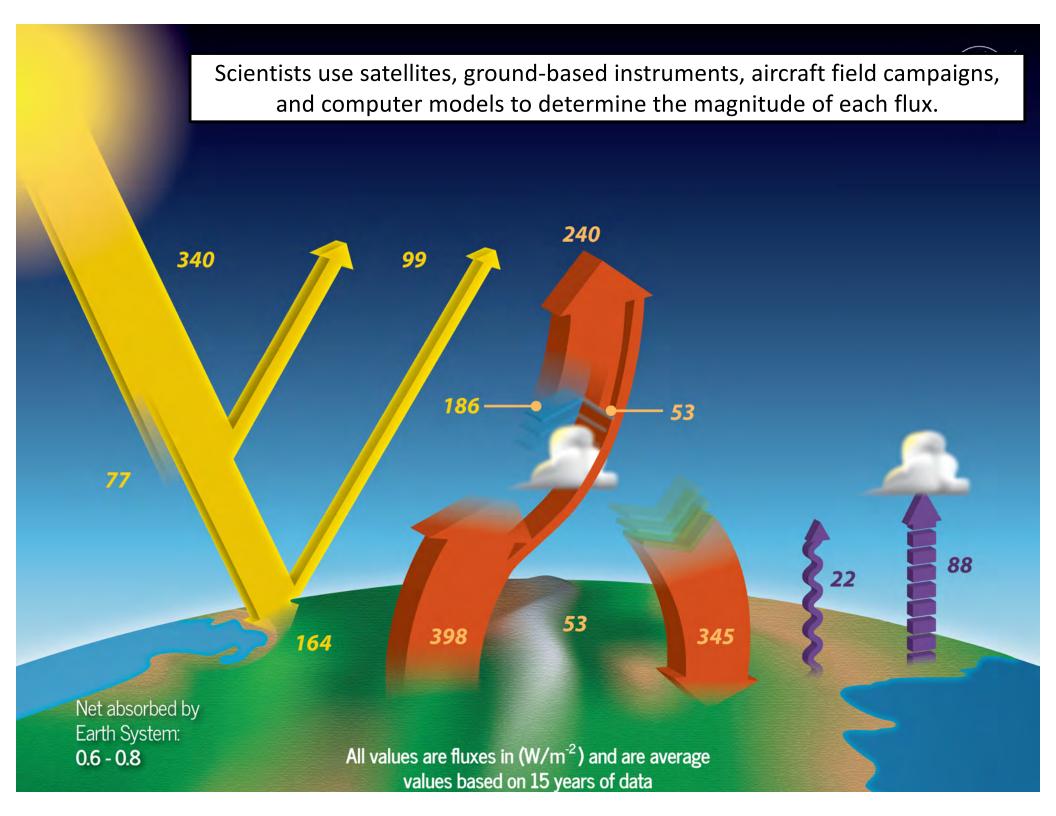






The world in emitted heat, May 25, 2001. Deserts are hot; clouds and polar ice are cold. The south pole is in winter deep freeze.



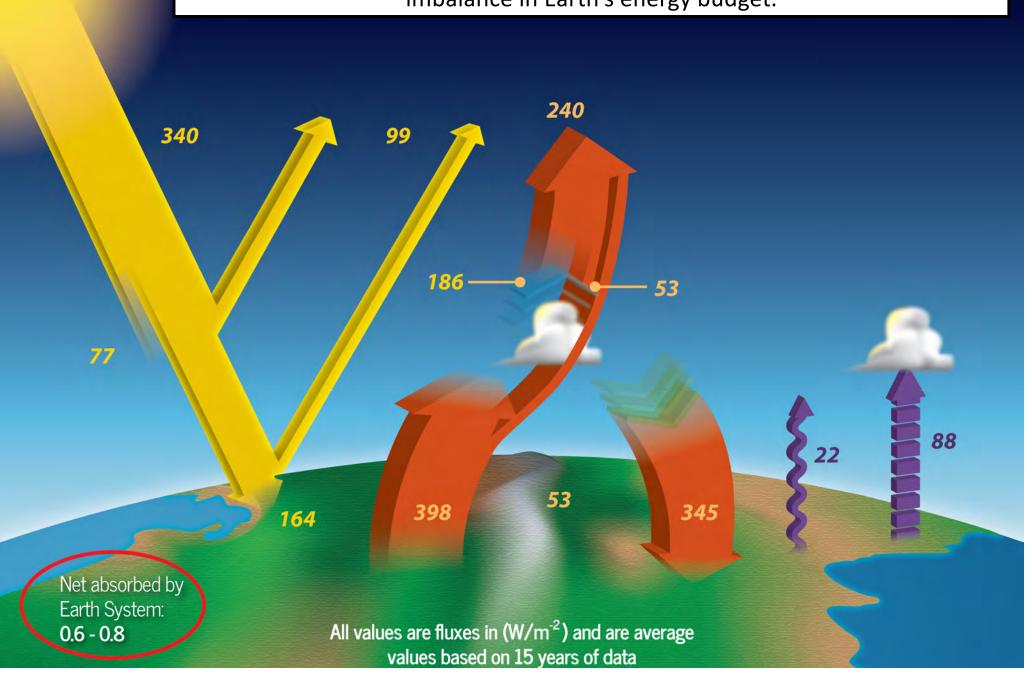




Like your house, anything that increases or decreases the amount of incoming or outgoing energy would disturb Earth's energy balance and would cause global temperatures to rise or fall.

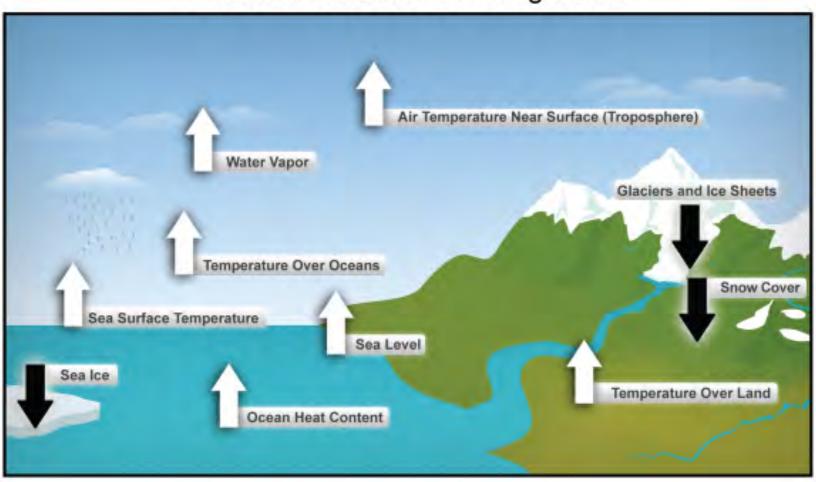


Over the last decade, our best estimate is that there is a small positive imbalance in Earth's energy budget.



This is consistent with several other *lines of evidence* of a warming planet.

Ten Indicators of a Warming World

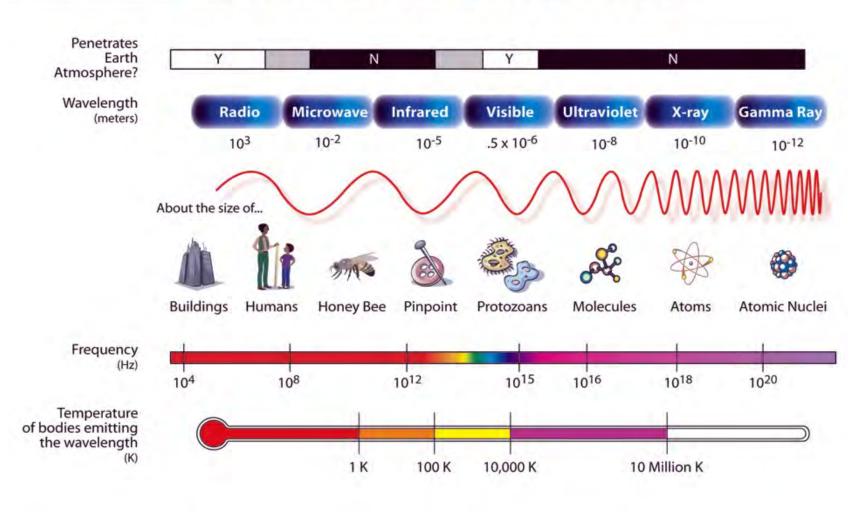


The End ... for now

Details behind the story

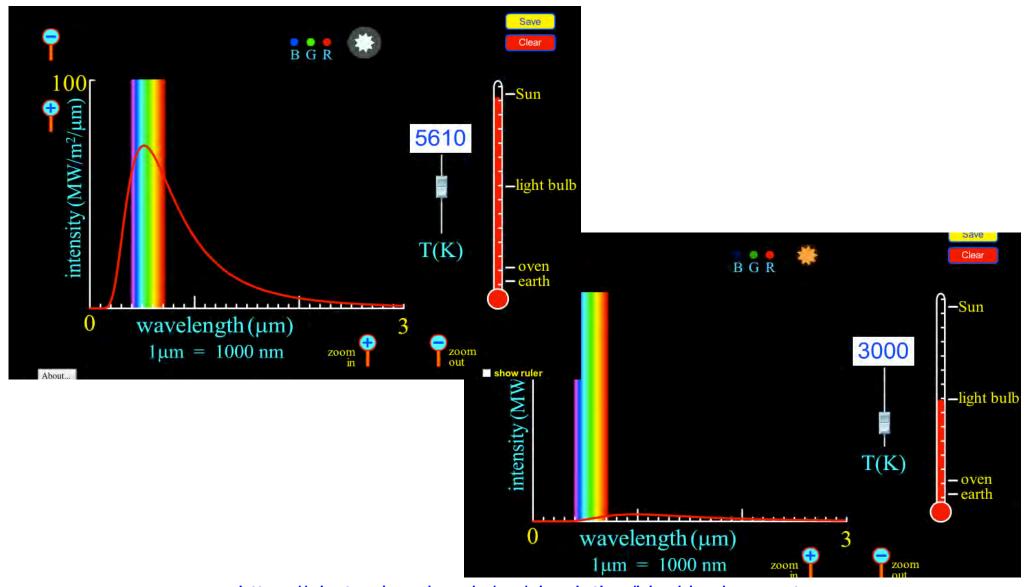
Teaching Resources: The Electromagnetic Spectrum

THE ELECTROMAGNETIC SPECTRUM



https://mynasadata.larc.nasa.gov/basic-page/electromagnetic-spectrum-diagram

The Blackbody Spectrum



https://phet.colorado.edu/en/simulation/blackbody-spectrum

Equil Temp Calculation – An Equation!

Equil Temp = Temp_{star} * (1-albedo)
$$^{1/4}$$
*

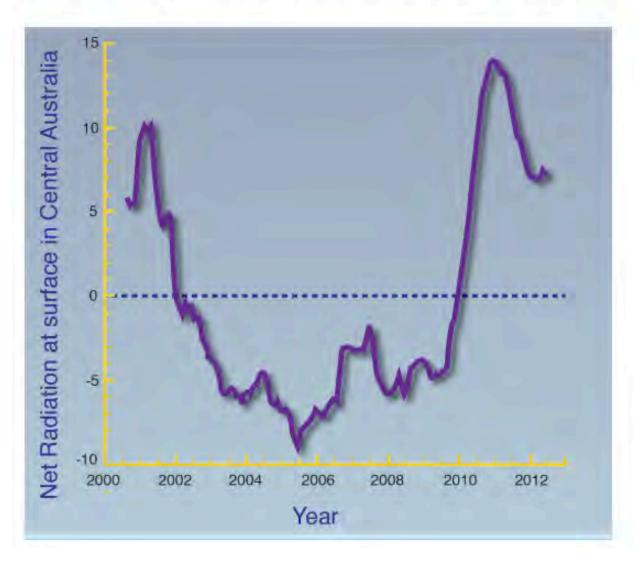
Square root
$$\left(\begin{array}{c} \text{Radius}_{\text{star}} \\ 2 & \text{Distance} \end{array}\right)$$

Temp _{Sun} ~ 5778K	Radius _{sun} ~ 695,500 km
Albedo ~ 0.3	Distance ~ 149,600,000 km

For Earth: Equil Temp ~254 K = -18 Celsius

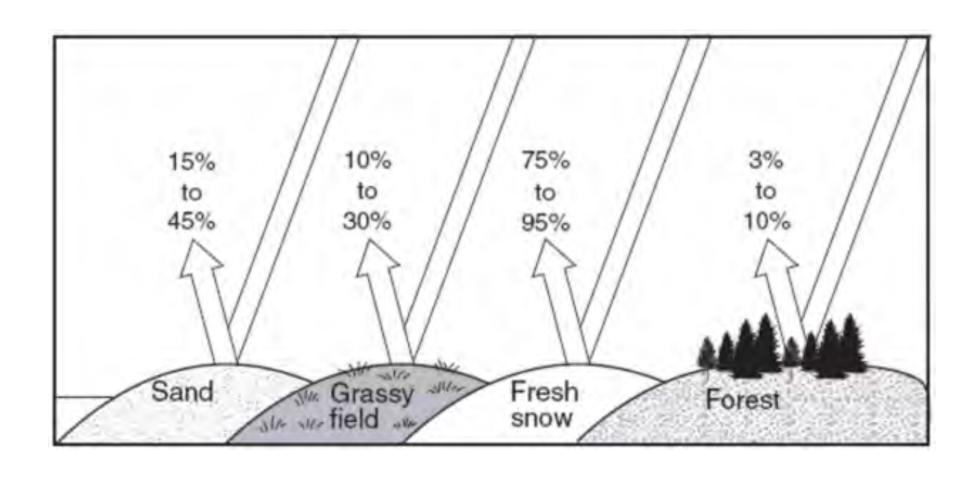
Energy Budget Mini Lesson

CASE STUDY: Surface Effects on Energy Budget - Changes in Central Australia



https://mynasadata.larc.nasa.gov/maps-graphs-and-data/earths-energy-budget

Curriculum Unit Plan



Unit: Earth's Energy Budget

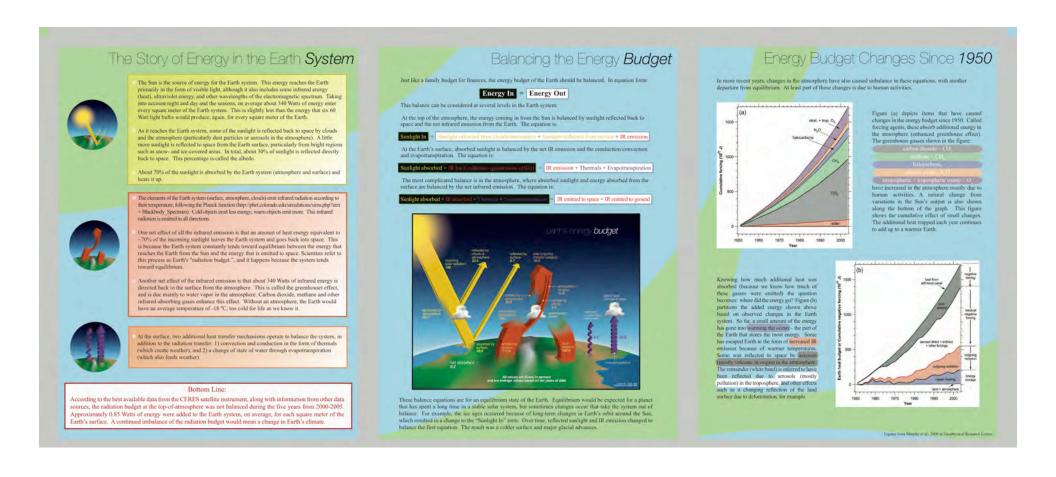
Greenhouse Effect



https://www.youtube.com/watch?v=b3NFfWlfj24

Teaching Resources

http://science-edu.larc.nasa.gov/energy_budget/



More Background Resources

https://science-edu.larc.nasa.gov/energy_budget/links.html



Teaching Resources: Explore Data

Earth System Data Explorer

